

A GEOMETRIC INEQUALITY ON HYPERSURFACE IN HYPERBOLIC SPACE

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ABSTRACT. In this paper, we use the inverse curvature flow to prove a sharp geometric inequality on star-shaped and two-convex hypersurface in hyperbolic space.

1. INTRODUCTION

The classical Alexandrov-Fenchel inequalities for closed convex hypersurface $\Sigma \subset \mathbb{R}^n$ state that

$$\int_{\Sigma} \sigma_m(\kappa) d\mu \geq C_{n,m} \left(\int_{\Sigma} \sigma_{m-1}(\kappa) d\mu \right)^{\frac{n-m-1}{n-m}}, \quad 1 \leq m \leq n-1 \quad (1)$$

where $\sigma_m(\kappa)$ is the m -th elementary symmetric polynomial of the principal curvatures $\kappa = (\kappa_1, \dots, \kappa_{n-1})$ of Σ and $C_{n,m} = \frac{\sigma_m(1, \dots, 1)}{\sigma_{m-1}(1, \dots, 1)}$ is a constant. When $m = 0$, (1) is interpreted as the classical isoperimetric inequality

$$|\Sigma|^{\frac{1}{n-1}} \geq \bar{C}_n Vol(\Omega)^{\frac{1}{n}}, \quad (2)$$

which holds on all bounded domain $\Omega \subset \mathbb{R}^n$ with boundary $\Sigma = \partial\Omega$. Here $|\Sigma|$ is the area Σ and \bar{C}_n is a constant depending only on dimension n . Inequality (1) was generalized to star-shaped and m -convex hypersurface $\Sigma \subset \mathbb{R}^n$ by Guan and Li [8] using the inverse curvature flow recently, where m -convex means that the principal curvature of Σ lies in the Garding's cone

$$\Gamma_m = \{\kappa \in \mathbb{R}^{n-1} | \sigma_i(\kappa) > 0, i = 1, \dots, m\}.$$

Recently, Huisken [11] showed that in the case $m = 1$, the assumption *star-shaped* can be replaced by *outward-minimizing*.

In this paper, we consider the hyperbolic space $\mathbb{H}^n = \mathbb{R}^+ \times \mathbb{S}^{n-1}$ endowed with the metric

$$\bar{g} = dr^2 + \sinh^2 r g_{\mathbb{S}^{n-1}},$$

where $g_{\mathbb{S}^{n-1}}$ is the standard round metric on the unit sphere \mathbb{S}^{n-1} . It's a natural question to establish some analogue inequalities of (1) for closed hypersurface in \mathbb{H}^n . In the case of $m = 1$, $\sigma_1 = \sigma_1(\kappa)$ is just the mean curvature H of Σ . Gallego and Solanes [6] have obtained a generalization of (1) to convex hypersurface in hyperbolic space using integral geometric methods,

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however, their result does not seem to be sharp. Denoting $\lambda(r) = \sinh r$, then $\lambda'(r) = \cosh r$. Recently, Brendle, Hung and Wang [3] proved the following inequality for star-shaped and mean convex (i.e., $H > 0$) hypersurface $\Sigma \subset \mathbb{H}^n$:

$$\int_{\Sigma} (\lambda' H - (n-1) \langle \bar{\nabla} \lambda', \nu \rangle) d\mu \geq (n-1) \omega_{n-1}^{\frac{1}{n-1}} |\Sigma|^{\frac{n-2}{n-1}} \quad (3)$$

where $|\Sigma|$ is the area of Σ and ω_{n-1} is the area of the unit sphere $\mathbb{S}^{n-1} \subset \mathbb{R}^n$. de Lima and Girao [4] also proved the following related inequality independently.

$$\int_{\Sigma} \lambda' H d\mu \geq (n-1) \omega_{n-1} \left(\left(\frac{|\Sigma|}{\omega_{n-1}} \right)^{\frac{n-2}{n-1}} + \left(\frac{|\Sigma|}{\omega_{n-1}} \right)^{\frac{n}{n-1}} \right), \quad (4)$$

Both inequalities (3) and (4) are sharp in the sense of that equality holds if and only if Σ is a geodesic sphere centered at the origin. Here, we say a closed hypersurface $\Sigma \subset \mathbb{H}^n$ is star-shaped if the unit outward normal ν satisfies $\langle \nu, \partial_r \rangle > 0$ everywhere on Σ , which is also equivalent to that Σ can be parametrized by a graph

$$\Sigma = \{(r(\theta), \theta) | \theta \in \mathbb{S}^{n-1}\}$$

for some smooth function r on \mathbb{S}^{n-1} . We note that inequalities (3) and (4) have some applications in general relativity, see [3, 4, 14].

In this paper, we consider the case $m = 2$. We prove the following sharp inequality for star-shaped and two-convex hypersurface $\Sigma \subset \mathbb{H}^n$, where *two-convex* means that the principal curvature lies in the Garding's cone Γ_2 everywhere on Σ .

Theorem 1. *If $\Sigma \subset \mathbb{H}^n$ is a star-shaped and two-convex hypersurface, then*

$$\int_{\Sigma} \sigma_2 d\mu \geq \frac{(n-1)(n-2)}{2} \left(\omega_{n-1}^{\frac{2}{n-1}} |\Sigma|^{\frac{n-3}{n-1}} + |\Sigma| \right), \quad (5)$$

where ω_{n-1} is the area of the unit sphere $\mathbb{S}^{n-1} \subset \mathbb{R}^n$ and $|\Sigma|$ is the area of Σ . The equality holds if and only if Σ is a geodesic sphere.

Note that there exists at least one elliptic point on a closed, connected hypersurface Σ in hyperbolic space \mathbb{H}^n . Proposition 3.2 in [1] shows that if σ_2 is positive, then σ_1 is automatically positive. So our assumption *two-convex* can also be replaced by $\sigma_2 > 0$ on Σ .

The proof of Theorem 1 follows a similar argument as in [3, 4, 8]. We evolve Σ by a special case of the inverse curvature flow in [7], and consider the following quantity defined by

$$Q(t) = |\Sigma|^{-\frac{n-3}{n-1}} \left(\int_{\Sigma} \sigma_2 d\mu - \frac{(n-1)(n-2)}{2} |\Sigma| \right).$$

We show that $Q(t)$ is monotone decreasing under the flow. Then we use the convergence result of the flow proved by Gerhardts to estimate a lower

bound of the limit of $Q(t)$:

$$\liminf_{t \rightarrow \infty} Q(t) \geq \frac{(n-1)(n-2)}{2} \omega_{n-1}^{\frac{2}{n-1}}.$$

In order to estimate this \liminf , we also use a sharp version Sobolev inequality on \mathbb{S}^{n-1} due to Beckner [2] as in [3]. Finally Theorem 1 follows easily from the monotonicity and the lower bound of $\liminf_{t \rightarrow \infty} Q(t)$.

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2. PRELIMINARIES

Let $\Sigma \subset \mathbb{H}^n$ be a closed hypersurface with unit outward normal ν . The second fundamental form h of Σ is defined by

$$h(X, Y) = \langle \bar{\nabla}_X \nu, Y \rangle$$

for any two tangent fields X, Y . The principal curvature $\kappa = (\kappa_1, \dots, \kappa_{n-1})$ are the eigenvalues of h with respect to the induced metric g on Σ . For $1 \leq m \leq n-1$, the m -th elementary symmetric polynomial of κ is defined as

$$\sigma_m(\kappa) = \sum_{i_1 < i_2 < \dots < i_m} \kappa_{i_1} \cdots \kappa_{i_m},$$

which can also be viewed as function of the second fundamental form $h_i^j = g^{jk} h_{ki}$. In the sequel, we will simply write σ_m for $\sigma_m(\kappa)$. We first collect the following basic facts on σ_m (see, e.g., [9, 12, 13]):

Lemma 2. Denote $(T_{m-1})_j^i = \frac{\partial \sigma_m}{\partial h_i^j}$ and $(h^2)_i^j = g^{jl} g^{pk} h_{kl} h_{ip}$. We have

$$\sum_{i,j} (T_{m-1})_j^i h_i^j = m \sigma_m, \quad (6)$$

$$\sum_{i,j} (T_{m-1})_j^i \delta_i^j = (n-m) \sigma_{m-1} \quad (7)$$

$$\sum_{i,j} (T_{m-1})_j^i (h^2)_i^j = \sigma_1 \sigma_m - (m+1) \sigma_{m+1} \quad (8)$$

Moreover, if $\kappa \in \Gamma_m^+$, we have the following Newton-MacLaurin inequalities

$$\frac{\sigma_{m-1} \sigma_{m+1}}{\sigma_m^2} \leq \frac{m(n-m-1)}{(m+1)(n-m)} \quad (9)$$

$$\frac{\sigma_1 \sigma_{m-1}}{\sigma_m} \geq \frac{m(n-1)}{n-m}, \quad (10)$$

and the equalities hold in (9), (10) at a given point if and only if Σ is umbilical there.

We now evolve $\Sigma \subset \mathbb{H}^n$ by the following evolution equation

$$\partial_t X = F\nu, \quad (11)$$

where ν is the unit outward normal to $\Sigma_t = X(t, \cdot)$ and F is the speed function which may depend on the position vector, principal curvatures and time. Let g_{ij} be the induced metric and $d\mu_t$ be its area element on Σ_t . We have the following evolution equations.

Proposition 3. *Under the flow (11), we have:*

$$\begin{aligned} \partial_t g_{ij} &= 2Fh_{ij} \\ \partial_t \nu &= -\nabla F, \\ \partial_t h_i^j &= -\nabla^j \nabla_i F - F(h^2)_i^j + F\delta_i^j, \\ \partial_t d\mu &= F\sigma_1 d\mu, \end{aligned} \quad (12)$$

$$\begin{aligned} \partial_t \sigma_m &= -\nabla^i ((T_{m-1})_i^j \nabla_j F) - F(\sigma_1 \sigma_m - (m+1)\sigma_{m+1}) \\ &\quad + (n-m)F\sigma_{m-1}, \end{aligned} \quad (13)$$

Proof. The first four equations follow from direct computations like in [10]. Now we calculate the evolution of σ_m (cf. [8])

$$\begin{aligned} \partial_t \sigma_m &= \frac{\partial \sigma_m}{\partial h_i^j} \partial_t h_i^j \\ &= - (T_{m-1})_j^i \nabla^j \nabla_i F - F(T_{m-1})_j^i (h^2)_i^j + F(T_{m-1})_j^i \delta_i^j \\ &= - \nabla^j ((T_{m-1})_j^i \nabla_i F) - F(\sigma_1 \sigma_m - (m+1)\sigma_{m+1}) \\ &\quad + (n-m)F\sigma_{m-1}, \end{aligned}$$

where in the last equality we used (7),(8) and the divergence free property of $(T_{m-1})_j^i$ (see [13]). \square

Proposition 4. *Under the flow (11), we have*

$$\frac{d}{dt} \int_{\Sigma} \sigma_m d\mu = (m+1) \int_{\Sigma} F\sigma_{m+1} d\mu + (n-m) \int_{\Sigma} F\sigma_{m-1} d\mu.$$

Proof. This proposition follows directly from (12), (13) and the divergence theorem. \square

In [7] Gerhardts studied general inverse curvature flow of star-shaped hypersurface in hyperbolic space. For our purpose, we will use a special case of their result for the following flow

$$\partial_t X = \frac{n-2}{2(n-1)} \frac{\sigma_1}{\sigma_2} \nu. \quad (14)$$

Theorem 5 (Gerhardt [7]). *If the initial hypersurface is star-shaped and strictly two-convex, then the solution for the flow (14) exists for all time $t > 0$ and the flow hypersurfaces converge to infinity while maintaining star-shapedness and strictly two-convex. Moreover, the hypersurfaces become*

strictly convex exponentially fast and more and more totally umbilical in the sense of

$$|h_i^j - \delta_i^j| \leq Ce^{-\frac{t}{n-1}}, \quad t > 0,$$

i.e., the principal curvatures are uniformly bounded and converge exponentially fast to one.

3. MONOTONICITY

We define the quantity

$$Q(t) = |\Sigma_t|^{-\frac{n-3}{n-1}} \left(\int_{\Sigma_t} \sigma_2 d\mu - \frac{(n-1)(n-2)}{2} |\Sigma_t| \right),$$

where $|\Sigma_t|$ is the area of Σ_t . In this section, we show that $Q(t)$ is monotone decreasing along the flow (14).

Proposition 6. *Under the flow (14), the quantity $Q(t)$ is monotone decreasing. Moreover, $\frac{d}{dt}Q(t) = 0$ at some time t if and only if the surface Σ_t is totally umbilical.*

Proof. Under the flow (14), Proposition 4 and (12) imply that

$$\frac{d}{dt} \int_{\Sigma} \sigma_2 d\mu = \frac{3(n-2)}{2(n-1)} \int_{\Sigma} \frac{\sigma_1 \sigma_3}{\sigma_2} d\mu + \frac{(n-2)^2}{2(n-1)} \int_{\Sigma} \frac{\sigma_1^2}{\sigma_2} d\mu \quad (15)$$

$$\frac{d}{dt} |\Sigma_t| = \frac{(n-2)}{2(n-1)} \int_{\Sigma} \frac{\sigma_1^2}{\sigma_2} d\mu. \quad (16)$$

Combining (15), (16) and (9), we have

$$\frac{d}{dt} \left(\int_{\Sigma} \sigma_2 d\mu - (n-2) |\Sigma_t| \right) \leq \frac{n-3}{n-1} \int_{\Sigma} \sigma_2 d\mu. \quad (17)$$

By applying the Newton-MacLaurin inequality (10) in (16), we also have

$$\frac{d}{dt} |\Sigma_t| \geq |\Sigma_t|. \quad (18)$$

Then combining (17) and (18) gives that

$$\frac{d}{dt} \left(\int_{\Sigma} \sigma_2 d\mu - \frac{(n-1)(n-2)}{2} |\Sigma_t| \right) \leq \frac{n-3}{n-1} \left(\int_{\Sigma} \sigma_2 d\mu - \frac{(n-1)(n-2)}{2} |\Sigma_t| \right). \quad (19)$$

From Proposition 8 in the next section and (19), we know that the quantity

$$\int_{\Sigma} \sigma_2 d\mu - \frac{(n-1)(n-2)}{2} |\Sigma_t|$$

is nonnegative along the flow (14). Then inequalities (18) and (19) implies that

$$\frac{d}{dt} Q(t) \leq 0.$$

If the equality holds, the inequalities (9) and (10) assume equalities everywhere on Σ_t . Then Σ_t is totally umbilical. \square

4. THE ASYMPTOTIC BEHAVIOR OF MONOTONE QUANTITY

In this section, we use the convergence result of the flow (14) proved in [7] to estimate the lower bound of the limit of $Q(t)$. First we need the following sharp Sobolev inequality on \mathbb{S}^{n-1} ([2]).

Lemma 7. *For every positive function f on \mathbb{S}^{n-1} , we have*

$$\begin{aligned} \int_{\mathbb{S}^{n-1}} f^{n-3} d\text{vol}_{\mathbb{S}^{n-1}} + \frac{n-3}{n-1} \int_{\mathbb{S}^{n-1}} f^{n-5} |\nabla f|^2 d\text{vol}_{\mathbb{S}^{n-1}} \\ \geq \omega_{n-1}^{\frac{2}{n-1}} \left(\int_{\mathbb{S}^{n-1}} f^{n-1} d\text{vol}_{\mathbb{S}^{n-1}} \right)^{\frac{n-3}{n-1}}. \end{aligned}$$

Moreover, equality holds if and only if f is a constant.

Proof. From Theorem 4 in [2], for any positive smooth function w on \mathbb{S}^{n-1} , we have the following inequality

$$\begin{aligned} \frac{4}{(n-1)(n-3)} \int_{\mathbb{S}^{n-1}} |\nabla w|^2 d\text{vol}_{\mathbb{S}^{n-1}} + \int_{\mathbb{S}^{n-1}} w^2 d\text{vol}_{\mathbb{S}^{n-1}} \\ \geq \omega_{n-1}^{\frac{2}{n-1}} \left(\int_{\mathbb{S}^{n-1}} w^{\frac{2(n-1)}{n-3}} d\text{vol}_{\mathbb{S}^{n-1}} \right)^{\frac{n-3}{n-1}}. \end{aligned}$$

Moreover equality holds if and only if w is constant. For any positive function f on \mathbb{S}^{n-1} , by letting $w = f^{\frac{n-3}{2}}$, we have

$$\begin{aligned} \int_{\mathbb{S}^{n-1}} f^{n-3} d\text{vol}_{\mathbb{S}^{n-1}} + \frac{n-3}{n-1} \int_{\mathbb{S}^{n-1}} f^{n-5} |\nabla f|^2 d\text{vol}_{\mathbb{S}^{n-1}} \\ \geq \omega_{n-1}^{\frac{2}{n-1}} \left(\int_{\mathbb{S}^{n-1}} f^{n-1} d\text{vol}_{\mathbb{S}^{n-1}} \right)^{\frac{n-3}{n-1}} \end{aligned}$$

and equality holds if and only if f is a constant. \square

Proposition 8. *Under the flow (14), we have*

$$\liminf_{t \rightarrow \infty} Q(t) \geq \frac{(n-1)(n-2)}{2} \omega_{n-1}^{\frac{2}{n-1}}. \quad (20)$$

Proof. Recall that star-shaped hypersurfaces can be written as graphs of function $r = r(t, \theta)$, $\theta \in \mathbb{S}^{n-1}$. Denote $\lambda(r) = \sinh(r)$, then $\lambda'(r) = \cosh(r)$. We next define a function φ on \mathbb{S}^{n-1} by $\varphi(\theta) = \Phi(r(\theta))$, where $\Phi(r)$ is a positive function satisfying $\Phi' = 1/\lambda$. Let $\theta = \{\theta^j\}$, $j = 1, \dots, n-1$ be a coordinate system on \mathbb{S}^{n-1} and φ_i, φ_{ij} be the covariant derivatives of φ with respect to the metric $g_{\mathbb{S}^{n-1}}$. Define

$$v = \sqrt{1 + |\nabla \varphi|_{\mathbb{S}^{n-1}}^2}.$$

From [7], we know that

$$\lambda = O(e^{\frac{t}{n-1}}), \quad |\nabla \varphi|_{\mathbb{S}^{n-1}} + |\nabla^2 \varphi|_{\mathbb{S}^{n-1}} = O(e^{-\frac{t}{n-1}}) \quad (21)$$

Since $\lambda' = \sqrt{1 + \lambda^2}$, we have

$$\lambda' = \lambda(1 + \frac{1}{2}\lambda^{-2} + O(e^{-\frac{4t}{n-1}})) \quad (22)$$

From (21) we also have

$$\frac{1}{v} = 1 - \frac{1}{2}|\nabla\varphi|_{\mathbb{S}^{n-1}}^2 + O(e^{-\frac{4t}{n-1}}) \quad (23)$$

In terms of φ , we can express the metric and second fundamental form of Σ as following (see, e.g, [3, 5])

$$\begin{aligned} g_{ij} &= \lambda^2(\sigma_{ij} + \varphi_i\varphi_j) \\ h_{ij} &= \frac{\lambda'}{v\lambda}g_{ij} - \frac{\lambda}{v}\varphi_{ij}, \end{aligned}$$

where $\sigma_{ij} = g_{\mathbb{S}^{n-1}}(\partial_{\theta^i}, \partial_{\theta^j})$. Denote $a_i = \sum_k \sigma^{ik}\varphi_{ki}$ and note that $\sum_i a_i = \Delta_{\mathbb{S}^{n-1}}\varphi$. By (21), the principal curvatures of Σ_t has the following form

$$\kappa_i = \frac{\lambda'}{v\lambda} - \frac{a_i}{v\lambda} + O(e^{-\frac{4t}{n-1}}), \quad i = 1, \dots, n-1.$$

Then we have

$$\begin{aligned} \sigma_2 &= \sum_{i < j} \kappa_i \kappa_j \\ &= \frac{(n-1)(n-2)}{2} \left(\frac{\lambda'}{v\lambda} \right)^2 - (n-2) \frac{\lambda' \Delta_{\mathbb{S}^{n-1}}\varphi}{v^2 \lambda^2} + O(e^{-\frac{4t}{n-1}}). \end{aligned}$$

By using (22) and (23),

$$\begin{aligned} \sigma_2 &= \frac{(n-1)(n-2)}{2} (1 + \lambda^{-2} - |\nabla\varphi|_{\mathbb{S}^{n-1}}^2) \\ &\quad - (n-2)\lambda^{-1} \Delta_{\mathbb{S}^{n-1}}\varphi + O(e^{-\frac{4t}{n-1}}). \end{aligned}$$

On the other hand,

$$\sqrt{\det g} = (\lambda^{n-3} + O(e^{-\frac{(n-3)t}{n-1}})) \sqrt{\det g_{\mathbb{S}^{n-1}}}.$$

So we have

$$\begin{aligned}
& \int_{\Sigma_t} \sigma_2 d\mu - \frac{(n-1)(n-2)}{2} |\Sigma_t| \\
&= \int_{\mathbb{S}^{n-1}} \lambda^{n-1} \left(\sigma_2 - \frac{(n-1)(n-2)}{2} \right) d\text{vol}_{\mathbb{S}^{n-1}} + O(e^{\frac{(n-5)t}{n-1}}) \\
&= \frac{(n-1)(n-2)}{2} \int_{\mathbb{S}^{n-1}} (\lambda^{n-3} - \lambda^{n-1} |\nabla \varphi|_{\mathbb{S}^{n-1}}^2) d\text{vol}_{\mathbb{S}^{n-1}} \\
&\quad - (n-2) \int_{\mathbb{S}^{n-1}} \lambda^{n-2} \Delta_{\mathbb{S}^{n-1}} \varphi d\text{vol}_{\mathbb{S}^{n-1}} + O(e^{\frac{(n-5)t}{n-1}}) \\
&= \frac{(n-1)(n-2)}{2} \int_{\mathbb{S}^{n-1}} (\lambda^{n-3} - \lambda^{n-1} |\nabla \varphi|_{\mathbb{S}^{n-1}}^2) d\text{vol}_{\mathbb{S}^{n-1}} \\
&\quad + (n-2)^2 \int_{\mathbb{S}^{n-1}} \lambda^{n-3} \nabla \lambda \nabla \varphi d\text{vol}_{\mathbb{S}^{n-1}} + O(e^{\frac{(n-5)t}{n-1}}).
\end{aligned}$$

Since $\nabla \lambda = \lambda \lambda' \nabla \varphi$, by using (22), we deduce that

$$\begin{aligned}
& \int_{\Sigma_t} \sigma_2 d\mu - \frac{(n-1)(n-2)}{2} |\Sigma_t| \\
&= \frac{(n-1)(n-2)}{2} \int_{\mathbb{S}^{n-1}} \left(\lambda^{n-3} + \frac{n-3}{n-1} \lambda^{n-5} |\nabla \lambda|^2 \right) d\text{vol}_{\mathbb{S}^{n-1}} + O(e^{\frac{(n-5)t}{n-1}}).
\end{aligned} \tag{24}$$

Moreover,

$$|\Sigma_t|^{\frac{n-3}{n-1}} = \left(\int_{\mathbb{S}^{n-1}} \lambda^{n-1} d\text{vol}_{\mathbb{S}^{n-1}} \right)^{\frac{n-3}{n-1}} + O(e^{\frac{(n-5)t}{n-1}}). \tag{25}$$

Using Lemma 7, we can complete the proof of Proposition 8 by combining (24) and (25). \square

We now complete the proof of Theorem 1

Proof of Theorem 1. Since $Q(t)$ is monotone decreasing, we have

$$Q(0) \geq \liminf_{t \rightarrow \infty} Q(t) \geq \frac{(n-1)(n-2)}{2} \omega_{n-1}^{\frac{2}{n-1}}.$$

This gives that the initial hypersurface Σ satisfies

$$\left(\int_{\Sigma} \sigma_2 d\mu - \frac{(n-1)(n-2)}{2} |\Sigma| \right) \geq \frac{(n-1)(n-2)}{2} \omega_{n-1}^{\frac{2}{n-1}} |\Sigma|^{\frac{n-3}{n-1}},$$

which is equivalent to the inequality (5) in Theorem 1. Now we assume that equality holds in (5), which implies that $Q(t)$ is a constant. Then Proposition 6 implies Σ_t is umbilical and therefore a geodesic sphere. It is also easy to see that if Σ is a geodesic sphere of radius r , then the area of Σ

is $|\Sigma| = \omega_{n-1} \sinh^{n-1} r$ and the integral of σ_2 is

$$\begin{aligned} \int_{\Sigma} \sigma_2 d\mu &= \frac{(n-1)(n-2)}{2} \omega_{n-1} \coth^2 r \sinh^{n-1} r \\ &= \frac{(n-1)(n-2)}{2} \omega_{n-1} (\sinh^{n-1} r + \sinh^{n-3} r) \\ &= \frac{(n-1)(n-2)}{2} \left(|\Sigma| + \omega_{n-1}^{\frac{2}{n-1}} |\Sigma|^{\frac{n-3}{n-1}} \right). \end{aligned}$$

Hence the equality holds in (5) on a geodesic sphere. This completes the proof of Theorem 1. \square

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